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A novel approach for edge detection based on the theory of universal gravity

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Abstract

This paper presents a new, simple and effective low-level processing edge detection algorithm based on the law of universal gravity. The algorithm assumes that each image pixel is a celestial body with a mass represented by its grayscale intensity. Accordingly, each celestial body exerts forces onto its neighboring pixels and in return receives forces from the neighboring pixels. These forces can be calculated by the law of universal gravity. The vector sums of all gravitational forces along, respectively, the horizontal and the vertical directions are used to compute the magnitude and the direction of signal variations. Edges are characterized by high magnitude of gravitational forces along a particular direction and can therefore be detected. The proposed algorithm was tested and compared with conventional methods such as Sobel, LOG, and Canny using several standard images, with and without the contamination of Gaussian white noise and salt & pepper noise. Results show that the proposed edge detector is more robust under noisy conditions. Furthermore, the edge detector can be tuned to work at any desired scale. © 2007 Pattern Recognition Society. Published by Elsevier Ltd. All rights reserved.

Keywords: Edge detection; Image processing; The law of universal gravity

1. Introduction

Edges correspond to sharp variations of image intensity and convey vitally important information in an image. Detection of edges is therefore a key issue in image processing, computer vision, and pattern recognition. A variety of algorithms exist for edge characterization and detection such as the statistical methods [1-5], the difference methods [6-8], and curve fitting [9-13].

As the complexity of image edge detection, every algorithm has its advantages and disadvantages. For example, Smith and Brady [14] proposed a new approach for low-level feature extraction—SUSAN operator. It works on the principle of "Univalue Segments" and tries to provide robust signatures for each edge point. However, the characteristics (e.g., size) of this "Smallest Univalue Segment Assimilating Nucleus" are strongly influenced by the presence of edges and corners. On the other hand, noise contamination is always a problem and edge detection in noisy environment can be treated as an optimal linear filter design problem [15–19]. Canny [16] described what has since become one of the most widely used edge finding algorithms. The first step taken is the definition of criteria which an edge detector must satisfy. Based on these Canny defined an optimal filter, which can be efficiently approximated by the first derivative of Gaussian function in the 1-D case. Canny's filter was further extended to recursive filters [20], which provide a more efficient way for image noise filtering and edge detection.

In the last decade, there have been renewed interests in the wavelet theory, and applications in filtering, classification, and compression [21]. Wavelet and its associated multi-resolution analysis have also been applied for the characterization of image intensity variations. Mallat et al. [22] presented their wavelet domain multi-scale edge detection approaches. In their researches, the edges are classified as the singularity points that can be detected as the local maxima of gradient moduli or the zero-crossings of wavelet coefficients. In Ref. [23], the zero-crossings of M-band wavelet coefficients are located and viewed as the edges. These multi-scale edge detection approaches have made a significant improvement for the image edge detection. More recently, a new approach based on the

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discrete singular convolution (DSC) edge detection algorithm has been proposed [24]. However, up to now, it is still a challenge issue to develop robust edge detection algorithm. In this paper, a new edge detection algorithm is presented.

The method is based on the law of universal gravity. Every image point may be assumed as a celestial body, which has relationships with other neighboring image points. The proposed edge detector algorithm includes three steps. At first, the gravitational forces of a pixel exert on every other pixels around it are computed using the law of universal gravity. Secondly, the vector sum of all gravitational forces is calculated. At last, the magnitude and direction of the vector are used to detect image edges. Experiments indicate that the new approach is effective for edge detection especially under noisy conditions.

The rest of the paper is organized as the following: the law of universal gravity is briefly reviewed in Section 2. Then, the algorithm of the proposed edge detector is presented in Section 3. Furthermore, applications of the presented algorithm are given in Section 4. The performance is illustrated using a number of real images. Both noise free and noisy contaminated images are used for the experiments. Finally, conclusions are presented in Section 5.

2. Background on the law of universal gravity

Between any two objects that have mass, there exist attractive gravitational forces acting on each object separately. The two forces have the same magnitude with opposite directions as illustrated in Fig. 1.

According to Newton's law of universal gravitation [25], every object in the Universe attracts every other object with a force directed along the line of centers for the two objects that is proportional to the product of their masses and inversely proportional to the square of the separation between the two objects. Newton's law of universal gravitation can be written as a vector equation to account for the direction of the gravitational force as well as its magnitude. As illustrated in Fig. 1, the formulation states:

$$\vec{f}_{1,2} = \frac{Gm_1m_2\hat{r}_{2,1}}{\|\vec{r}_{2,1}\|^2} = \frac{Gm_1m_2\vec{r}_{2,1}}{\|\vec{r}_{2,1}\|^3},\tag{1}$$

where $\vec{f}_{1,2}$ is the force on object 1 due to object 2, *G* is the gravitational constant, m_1 and m_2 are the masses of the objects 1 and 2, respectively, \vec{r}_1 and \vec{r}_2 are the vector positions of the two objects, respectively, $r_{2,1} = ||\vec{r}_2 - \vec{r}_1||$ is the distance between objects 1 and 2, $\hat{r}_{2,1} = \vec{r}_2 - \vec{r}_1/||\vec{r}_2 - \vec{r}_1||$ is the unit vector from object 2 to 1.



Fig. 1. Newton's law of universal gravitation.

Likewise, the vector gravitational force exerted by object 1 on object 2 takes the form

$$\vec{f}_{2,1} = -\vec{f}_{1,2}.$$
(2)

3. Gravity edge detector

3.1. Illustration of the method

To construct an edge detector, we assume that every image point is a celestial body, which has some relationship with other image points within its neighborhood through gravitational forces. For points beyond a pre-specified range, we assume all gravitational forces are zero. For each image point, the magnitude and the direction of the vector sum of all gravitational forces the point exerts on its neighborhood conveys the vitally important information about an edge structure including the magnitude and the direction. The idea is inspired by USAN [14] concept which was introduced by Smith and Brady for feature extraction.

There are a number of definitions of an edge, each being applicable in various specific circumstances. One of the most common and most general definitions is the ideal step edge [26]. For notational simplicity we only describe the detection algorithm for vertically orientated edges under ideal step edge condition. The algorithm can be easily extended to horizontal and diagonal orientations. Edge structures in 90° directions are depicted in Fig. 2. Here we assume that the gray value of white pixels is I_1 , and that of black pixels I_2 , $I_2 > I_1$. F_1 represents the gravitational force between two white pixels per unit distance, F_2 between a white and a black, and F_3 between black and black, $F_3 > F_2 > F_1$ according to Eq. (1). The magnitude



Fig. 2. Basic edge structures oriented in 90° directions.

of the gravitational force of the center pixel exerts on each other pixel within its neighborhood is illustrated in Fig. 2, the arrow indicates the gravitational force direction, here $a = \frac{1}{2}$, for simplification, we only consider 3×3 neighborhood of four points, of which points 1 and 4 are latent edge points whereas points 2 and 3 are non-edge points.

As can be seen from Fig. 2, all image points have four double gravitational forces. For point 1, we first compute the magnitude sum of gravitational forces in both x and y directions, respectively,

$$F_x = \left(aF_2 * \frac{\sqrt{2}}{2} + F_1 + aF_1 * \frac{\sqrt{2}}{2}\right) - \left(aF_2 * \frac{\sqrt{2}}{2} + F_1 + aF_1 * \frac{\sqrt{2}}{2}\right) = 0,$$

$$F_y = \left(aF_2 * \frac{\sqrt{2}}{2} + F_2 + aF_2 * \frac{\sqrt{2}}{2}\right) - \left(aF_1 * \frac{\sqrt{2}}{2} + F_1 + aF_1 * \frac{\sqrt{2}}{2}\right) = \left(\frac{\sqrt{2}}{2} + 1\right) * (F_2 - F_1).$$

Then we calculate the magnitude and the direction of the vector sum of point 1 exerts on its neighborhood:

$$F^{1} = \sqrt{(F_{x})^{2} + (F_{y})^{2}} = \left(\frac{\sqrt{2}}{2} + 1\right)(F_{2} - F_{1}),$$

$$\theta^{1} = \arctan(F_{x}/F_{y}) = 0.$$

Using exactly the same method, we obtain the response of points 2–4:

$$F^{2} = F^{3} = 0,$$

$$\theta^{2} = \theta^{3} = \arctan(F_{x}/F_{y}) = \frac{\pi}{2},$$

$$F^{4} = (\frac{\sqrt{2}}{2} + 1)(F_{3} - F_{2}),$$

$$\theta^{4} = \arctan(F_{x}/F_{y}) = 0.$$

According to the above, the response F is 0 for non-edge points such as points 2 and 3, but for latent edge points such as points 1 and 4, F is greater than 0. In general, $F^1 \neq F^4$, we assume $F^4 > F^1 > 0 = F^2 = F^3$. Theoretically, we can set an appropriate threshold, for example F, and let $F^4 > F > F^1$. A pixel location is declared an edge location if the value of response exceeds the threshold, here point 4 is labeled as edge point. Performing the same steps at each image point an edge map is then created (e.g., line 4 is labeled as edge). The edgedirection information is given by θ .

When using the Sobel or Prewitt operators in Fig. 2, the pixels of Column 3 with point 1 and Column 4 with point 4 have the same response and are both labeled as edge. Therefore, our method is superior to Sobel and Prewitt operators according to the criteria "only one response to a single edge" which was described by Canny [16].

3.2. Generation of the edge operator

The entire algorithm of edge detection can be implemented as follows:

(1) For each image point g(i, j), we consider a $m \times n$ neighborhood Ω with pixels $x_{k,l} \in \Omega$ & $(k, l) \neq (i, j)$. For each point, the gravitational force of the point exerts on its neighboring pixels is computed using Eq. (1):

$$\vec{f}_{i,j;k,l} = \frac{Gm_{i,j}m_{k,l}\vec{r}}{\|\vec{r}\|^3},$$
(3)

where $f_{i,j;k,l}$ is the gravity g(k, l) exerts on g(i, j), $m_{i,j}$, $m_{k,l}$ are the grav values of the pixel g(i, j) and g(k, l), \vec{r} is the vector from g(k, l) to g(i, j), and

$$\|\vec{r}\| = \sqrt{(k-i)^2 + (l-j)^2}.$$
(4)

(2) Considering an image as 2D signals, in practical computing, we compute the gravitational forces in both x and y directions:

$$f_{i,j;k,l}^{x} = \vec{f}_{i,j;k,l} \sin \theta = \frac{Gm_{i,j}m_{k,l}r}{\|\vec{r}\|^{3}} \frac{k-i}{\vec{r}}$$

$$= \frac{Gm_{i,j}m_{k,l}(k-i)}{\|\vec{r}\|^{3}},$$

$$f_{i,j;k,l}^{y} = \vec{f}_{i,j;k,l} \cos \theta = \frac{Gm_{i,j}m_{k,l}\vec{r}}{\|\vec{r}\|^{3}} \frac{l-j}{\vec{r}}$$

$$= \frac{Gm_{i,j}m_{k,l}(l-j)}{\|\vec{r}\|^{3}}.$$
(5)

And the vector $\vec{f}_{i,j;k,l}$ can be calculated using the following formula:

$$\vec{f}_{i,j;k,l} = f_{i,j;k,l}^{x} \hat{x} + f_{i,j;k,l}^{y} \hat{y}.$$
(6)

The vector sum of all gravitational forces which the neighborhoods exert on g(i, j) is expressed as

$$\vec{F}_{i,j} = \sum \vec{f}_{i,j;k,l} = F^x \hat{x} + F^y \hat{y}$$

(k, l) $\in \Omega \& (k, l) \neq (i, j),$ (7)

where

$$F^{x} = \sum f_{i,j;k,l}^{x} \quad (k,l) \in \Omega \& (k,l) \neq (i,j),$$

$$F^{y} = \sum f_{i,j;k,l}^{y} \quad (k,l) \in \Omega \& (k,l) \neq (i,j).$$
(8)

(3) The edge strength response of point g(i, j) is produced by the magnitude of the vector $\vec{F}_{i,j}$, its edge direction is given by the direction of $\vec{F}_{i,j}$:

$$\|\vec{F}\| = \sqrt{(F_x)^2 + (F_y)^2},$$

$$\theta = \arctan(F_x/F_y),$$
(9)

where $\|\vec{F}\|$ is the magnitude of the vector sum $\vec{F}_{i,j}$, θ is the direction of the vector sum $\vec{F}_{i,j}$.

(4) Set an appropriate threshold to produce an edge map.

In fact, the effect of the algorithm is equal to a convolution mask to approximate the first derivative of the image brightness function. To illustrate this, substituting Eqs. (4) and (5) in Eq. (8) yields

$$F^{x} = \sum f_{i,j;k,l}^{x} = Gm_{i,j} \sum \frac{m_{k,l}(k-i)}{\left(\sqrt{(i-k)^{2} + (j-l)^{2}}\right)^{3}}$$

$$(k,l) \in \Omega \& (k,l) \neq (i,j),$$

$$F^{y} = \sum f_{i,j;k,l}^{y} = Gm_{i,j} \sum \frac{m_{k,l}(l-j)}{\left(\sqrt{(i-k)^{2} + (j-l)^{2}}\right)^{3}}$$

$$(k,l) \in \Omega \& (k,l) \neq (i,j),$$
(10)

Eq. (10) is a local computation, set the location g(i, j) as the original of the coordinate system, so

(i, j) = (0, 0)

and

$$\left(k = -\left\lfloor \frac{m}{2} \right\rfloor, \dots, \dots, \left\lfloor \frac{m}{2} \right\rfloor; \quad l = -\left\lfloor \frac{n}{2} \right\rfloor, \dots, \dots, \left\lfloor \frac{n}{2} \right\rfloor\right).$$

For convenience we replace $Gm_{i,j}$ with a constant *C* since it is constant in the convolving process. Eq. (10) can be simplified to

$$F^{x} = \sum f_{i,j;k,l}^{x} = C \sum \frac{m_{k,l}k}{(\sqrt{k^{2} + l^{2}})^{3}}$$

$$(k,l) \in \Omega \& (k,l) \neq (0,0),$$

$$F^{y} = \sum f_{i,j;k,l}^{y} = C \sum \frac{m_{k,l}l}{(\sqrt{k^{2} + l^{2}})^{3}}$$

$$(k,l) \in \Omega \& (k,l) \neq (0,0).$$
(11)

Eq. (11) equals to a convolving procedure.

Conventional operators, such as Prewitt and Sobel operators, are sensitivity to noise and tend to generate spurious edges. The effect caused by noise may be filtered out if more neighboring points are considered by the operator. Experimental results demonstrate that for noisy images an edge operator with a larger mask may provide a better edge result.

In practice, the *C* can be set to other value to act at some special circumstances instead of only being defined as the gray level of the center pixel of the mask. The parameters, *m* and *n*, can take different values. In this paper, we say mask = *x*. means m = n = x. For example, let C = 1, taking the mask size equal to 3, we obtain two 3×3 masks for calculating the first derivatives in the *x* and *y* directions, respectively:

$$\begin{bmatrix} \frac{-1}{2\sqrt{2}} & \frac{-1}{1\sqrt{1}} & \frac{-1}{2\sqrt{2}} \\ 0 & 0 & 0 \\ \frac{1}{2\sqrt{2}} & \frac{1}{1\sqrt{1}} & \frac{1}{2\sqrt{2}} \end{bmatrix} \text{ and } \begin{bmatrix} \frac{-1}{2\sqrt{2}} & 0 & \frac{1}{2\sqrt{2}} \\ \frac{-1}{1\sqrt{1}} & 0 & \frac{1}{1\sqrt{1}} \\ \frac{-1}{2\sqrt{2}} & 0 & \frac{-1}{2\sqrt{2}} \end{bmatrix}.$$

The parameter C can also take other values to achieve more effective results. We have found that C can convey the regional information around g(i, j) and makes the algorithm

more efficient under salt & pepper noise conditions when C = sig(g(i, j)). Here sig(g(i, j)) is a sigmoid function which is inspired by Etienne [27].

4. Experimental results

To demonstrate the efficiency of the proposed approach, we carried out computer experiments on gray-level images. We selected a few standard images which are either natural or human-made (Fig. 3). The Lena and the Couple images are both figure image. The Singapore airport image is a natural and non-textured image and the square is a synthetic image. The settings of these images vary from in-door scenes to out-door views. The resolution of all images is 8-bit per pixel. The square image is of the size of 256×256 pixels, the Singapore airport is of 719×905 , the rest two images are of 512×512 .

Most edge detection techniques utilize a post processing thresholding immediately after feature extraction to thin and extend edge contours. There are many well-established thresholding methods [28,29] and edge thinning techniques. In the present work, the edge detection consists of two steps: edge magnitude calculation and threshold. For magnitude result, we first get the contrast stretched image with brightness ranging from 0 to 255 by linear stretch method, then utilizes Otsu [30] method to generate a threshold. The procedure described here is also applied for the implementation of other standard edge detectors, which are used for comparison in the present study.

In the rest of this section, we conduct two groups of computer experiments to test the proposed approach. Group one is designed to investigate the performance of the present algorithm on the edge detection of real images, including clean and noise images. Group two is designed to objectively compare the performances of different edge detectors by using a



Fig. 3. A collection of sample images, where the square is of the size of 256×256 , the Singapore airport is of 719×905 , the rest are of 512×512 .



Fig. 4. Edge maps of the Lena image: (a) the present detector with mask = 3, C = 1 (b) the Sobel detector, and (c) the LOG detector.

Fig. 5. Edge maps of the Lena image: (a) the present detector with mask = 3, C = sig(g(i, j)), (b) the Sobel detector, and (c) the Canny detector.

computer generated image. They are, respectively, described in the following three subsections. A brief discussion is given in the last subsection.

4.1. Edge detection of clean images

It is well-known that both the Sobel and LOG detectors are the most commonly used edge operators as first and second order detection, respectively. For Lena image two edge maps obtained by the two detectors are shown in Fig. 4(b) and (c). An edge map obtained by the new detector is shown in Fig. 4(a) with parameters mask = 3, and C = 1. Apparently, the standard Sobel detector and gravity detector yield similar results. This is due to the fact that the Gravity detector uses the similar convolution mask as the Sobel to approximate the first derivative of the image brightness function. The edge map of LOG is slightly thinner, but it produces undesirable noise.

It is noted that none of the above-mentioned three edge detectors resolves some structures of the Lena image such as the nose feature. To illustrate the potential of the Gravity detector, we also conduct two tests by using Gravity with different parameters. The Gravity parameters are chosen as mask = 3, C = sig(g(i, j)). sig(g(i, j)) is the sigmoid function which has the form

$$sig(g(i, j)) = \frac{1}{1 + e^{\sigma(g(i, j) - \bar{g}(i, j))}},$$
(12)

where $\bar{g}(i, j)$ is the average intensity in the considered neighborhood at point g(i, j) and σ is the standard deviation in the neighborhood at the same point.

These results are depicted in Fig. 5, along with those obtained by Sobel and Canny detectors. Overall, the Canny detector produces smoother and thinner edges. However, it is clear that most of the edge structures are well preserved by the new detector. Although the edges by the present detector are slightly thicker than those of the Canny detector, its performance is still competitive with it, considering the Canny detector has a thinning stage of additional edge. In contrast, it is noted that the Gravity method detects more edge features (as the white ellipse indicated) than the Sobel operator. Furthermore, the edges detected by gravity operator are thinner than those of Sobel. Therefore, the gravity detector is more efficient than the Sobel detector for edge detection of these images.

4.2. Noisy images

To investigate the performance of the new algorithm under noisy environment, we consider a number of low grade images. We first compare the proposed detector with the Sobel and Canny detector for salt & pepper noise. The noise intensity is 0.1, which denotes significant degradation of image quality as show in Fig. 6.

For noisy Lena image, three edge maps obtained by these detectors are shown in Fig. 7: (a) the present detector with mask = 3, C = 1; (b) the Sobel detector; (c) the Canny detector. Obviously, the contrast of the Canny edge image is poor and contains much small and spurious contour. For comparison, much sharper edge images are successfully attained by the new detector as shown in Fig. 7(a). The better performance

Fig. 6. A noisy Lena and Couple image corrupted with salt & pepper noise with intensity = 0.1.

Fig. 7. Edge maps of the Lena image corrupted with additive salt & pepper noise (intensity = 0.1): (a) is obtained using the present detector with parameters mask = 3, C = 1; (b) and (c) are obtained using the Sobel and Canny detector, respectively.

1	1	0	-	1	_	1
1	1	0	-	1	-	1
1	1	0	-	1	-	1
1	1	0	-	1	-	1
_ 1	1	0	-	1	-	1

Fig. 8. The vertical mask of expanded Prewitt detector.

of the present detector in the presence of severe noise is also perceivable compared with Fig. 7(b), with much of the noise pulses removed and the edge map is more clear.

Now we expand the mask size to 5. The compared detectors are Prewitt and LOG detectors with the same mask size. Fig. 8 illustrates the vertical mask of expanded Prewitt detector [31]. Four edge maps of Couple image are show in Fig. 9. (a) and (b) are obtained using the present detector with parameters (a) mask = 5, C = 1; (b) mask = 5, C = sig(g(i, j)); (c) and (d) are obtained using the Prewitt and LOG detector with mask = 5, respectively. The better performance of the present detector in the presence of salt & pepper noise is clearly perceivable, compared to Fig.9(c), much of the noise pulses is removed and edge map is more lucid in Fig. 9(a).

This example also illustrates the effects of adjusting the parameters on the resultant edge map. Some of the fine edge structures are missing and the edges are slightly thicker in Fig. 9(c). Comparatively, the edges in Fig. 9(b) are thinner and most of the fine edges with relatively lower contrast are detected. Compared with Fig. 9(d), the edge map of Fig. 9(b) is also clearer. Hence, increasing mask size has a significant effect on the reduction of noise influence, and the parameter of C=sig(g(i, j)) results in thinner edges under salt & pepper noise conditions.

Results on the Singapore airport image corrupted with Gaussian noise (mean = 0, intensity = 0.1) and random noise is show in Fig. 10. Fig. 11 shows the resulting edge images detected from noisy environment, obtained by the present detector (Row 1), the expanded Prewitt detector (Row 2), the LOG detector (Row 3), and the Canny detector (Row 4). For all detectors, the mask size of Column 1 is 5 and the Column 2 is 7, respectively.

As can be seen from these edge maps, the impact of noise on these detectors is clearly visible: the LOG detector is ineffective under such noise conditions. In contrast, sharp edge images are successfully attained by using the proposed detector. There is a little visual difference between those obtained by the present detector and the expanded Prewitt detector, but with the mask expanding, the proposed algorithm detects more

Fig. 9. Edge maps of the Couple image corrupted with additive salt & pepper noise (intensity = 0.1): (a) and (b) are obtained using the present detector with parameters (a) C = 1, mask = 5 and (b), mask = 5; C = sig(g(i, j)); (c) and (d) are obtained using the expanded Prewitt and LOG detector, respectively, with mask = 5.

Fig. 10. A noisy Singapore airport image corrupted with Gaussian noise (mean = 0, intensity = 0.1) and random noise.

edges efficiently: when the mask size is expanded to 7, the proposed detector can detect more edges, including some fine structural edges (for instance the right bottom corner).

As mentioned in the introduction, the Canny detector [16] was formulated as an optimization problem for being used under noise environment. It is demonstrated in Refs. [32,33] that a smooth parameter $\sigma = 1.5$ is nearly optimal in association with a 5 × 5 mask. The resulting edge images are included in Row 4 of Fig. 11 for comparison. Understandably, with the mask expanding, the Canny operator can detect narrower edges. On the contrary, the proposed detector (Row 1) can detect wider edges therefore fine edges may be detected. It is demonstrated that the new detector with a larger mask may provide a wider edge map under such condition compared with others.

4.3. Edge detection of a synthetic image

To validate the Gravity detector further, we present an alternative evaluation in this subsection. Edge detection systems could be compared in many ways. For example, image gradients may be compared visually [33], where an edge image is evaluated by a group of people and the average score may serve as an index of quality. For synthetic images, where the exact location of edges is known, Abdou and Pratt [34] proposed a figure of merit to objectively evaluate the performance

Fig. 11. Edge maps of the Singapore airport corrupted with Gaussian noise (mean = 0, intensity = 0.1) and random noise: (Row 1) the present detector; (Row 2) the expanded Prewitt detector; (Row 3) the LOG detector, and (Row 4) the Canny detector. The mask size of Column 1 is 5 and Column 2 is 7.

of edge detectors. Their figure of merit is defined as

$$F = \frac{1}{\max(N_l, N_D)} \sum_{i=1}^{N_D} \frac{1}{1 + \xi d_i^2},$$
(13)

where d_i is the distance between a pixel declared as edge point and the nearest ideal edge pixel, ξ is a scaling constant set to $\frac{1}{9}$ as in Pratt's work. N_l and N_D are the numbers of ideal and detected edge pixels, respectively. It is common practice to evaluate the performance of an edge detector for synthetic images by introducing noise in the images. *F* is an index to measure the accuracy of edge localizations. A larger value corresponds to better performance, with 1 being a perfect result.

Performance comparison is based on a synthetic square image as shown in Fig. 3. The figure of merit *F* for each of the methods studied is calculated with respect to different power of the added Gaussian white noise, and the results are shown in Table 1. Here Gravity.1 with mask = 3, C = 1; Gravity.2 with mask = 5, C = sig(g(i, j)); Sobel with mask = 3; LOG and Canny are both with mask = 5, as Section 4.2. When the noise level is 0, the *F* values are close to 1 and the performances of all the tested detectors are very satisfactory. With the increase of

Table 1 The statistic of the merit F for different edge detectors

Power of	Detector						
lioise	Gravity.1	Gravity.2	Sobel	LOG	Canny		
0	0.8375	0.9054	0.8342	0.9054	0.9341		
0.05	0.8412	0.9366	0.8301	0.1776	0.2778		
0.1	0.8101	0.8608	0.7059	0.1774	0.2219		
0.15	0.6006	0.6438	0.3947	0.1768	0.3194		
0.2	0.3987	0.4710	0.2706	0.1763	0.2945		
0.25	0.288	0.3810	0.2341	0.1753	0.2940		

the noise level, the F values decrease. In contrast, the proposed detectors achieve large F values over the domain of interest, suggesting their superiority over the other three detectors.

5. Discussions

In the presence of noise, the direct application of the differentiation operation in edge detection will encounter difficulties, as illustrated by the preceding experiments. The differentiation operation is sensitive to noise and the problem is mathematically ill-posed. To offset the effect of noise, a direct approach is to remove noise before the differentiation, usually by convolving the raw input image with a Gaussian function, which leads to the well-known LOG detector [8]. This problem can also be solved by using regularization techniques developed for dealing with mathematically ill-posed problems [35]. Poggio et al. [36] proved that the varying formulation of Tikhonov regularization leads to a Gaussian-like convolution filter. In the present work, the impact caused by noise would be reduced by using a larger mask. What's more, with the mask expanding, larger edges can be detected.

On the other hand, the algorithm has an extra parameter C, which can be tuned to act at certain special circumstances. Experiments in Section 4.1 indicated that when C = sig(g(i, i)), more detailed edges can be detected under salt & pepper noise condition. Considering the combined use of mask scale and C, the Gravity algorithm is a potential method for more complicated tasks.

6. Conclusions

A new algorithm which is efficient for edge detection based on the law of universal gravity is proposed. The performance of the proposed algorithm is compared with many other existing methods, including the Sobel, LOG, and Canny detectors. Experiments on a variety of images have shown that the algorithm is consistent and reliable even when image quality is significantly degraded by noise, especially salt & pepper noise. The algorithm requires two parameters. A series of edge operators with various sizes of masks are obtained from m, n. By considering a larger neighborhood, the effect of noise on edges is reduced. With the same size mask, the proposed detector is more effective than others. The combined use of the parameter C makes the present algorithm efficient for edge detection in a variety of practical situations, especially in salt & pepper noise conditions.

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